



PRESS RELEASE | PARIS | 04 JULY 2016

A giant impact: solving the mystery of how Mars' moons formed

Where did the two natural satellites of Mars, Phobos and Deimos, come from? For a long time, their shape suggested that they were asteroids captured by Mars. However, the shape and course of their orbits contradict this hypothesis. Two independent and complementary studies provide an answer to this question. One of these studies, to be published in *The Astrophysical Journal* and predominantly conducted by researchers from the CNRS and Aix-Marseille Université¹, rules out the capture of asteroids, and shows that the only scenario compatible with the surface properties of Phobos and Deimos is that of a giant collision. In the second study, a team of French, Belgian, and Japanese researchers used cutting-edge digital simulations to show how these satellites were able to form from the debris of a gigantic collision between Mars and a protoplanet one-third its size. This research, which is the result of collaboration between researchers from Université Paris Diderot and Royal Observatory of Belgium, in collaboration with the CNRS, Université de Rennes 1² and the Japanese Institute ELSI, is published on July 4, 2016 in the journal *Nature Geoscience*.

The origin of the two Martian moons, Phobos and Deimos, remained a mystery. Due to their small size and irregular shape, they strongly resembled asteroids, but no one understood how Mars could have « captured » them and made them into satellites with almost circular and equatorial orbits. According to a competing theory, toward the end of its formation Mars suffered a giant collision with a protoplanet: but why did the debris from such an impact create two small satellites instead of one enormous moon, like the Earth's? A third possibility is that Phobos and Deimos formed at the same time as Mars, which would entail that they have the same composition as their planet, although their low density seems to contradict this hypothesis. Two independent studies have now solved the puzzle: the Martian moons must have arisen from a giant collision.

In one of these studies, a team of Belgian, French, and Japanese researchers offers, for the first time, a complete and coherent scenario for the formation of Phobos and Deimos, which would have been created following a collision between Mars and a primordial body one-third its size, 100 to 800 million years after the beginning of the planet's formation. According to researchers, the debris from this collision formed a very wide disk around Mars, made up of a dense inner part composed of matter in fusion, and a very thin

¹ French laboratories involved: Laboratoire d'astrophysique de Marseille (CNRS/Aix-Marseille Université), Institut de planétologie et d'astrophysique de Grenoble (CNRS/Université Grenoble Alpes), Centre européen de recherche et d'enseignement de géosciences de l'environnement (CNRS/Aix-Marseille Université/IRD/Collège de France). ESA's European Space Research and Technology Centre was also involved.

² French laboratories involved: Institut de physique du globe de Paris (CNRS/IPGP/Université Paris Diderot), Institut de physique de Rennes (CNRS/Université de Rennes 1).



outer part primarily of gas. In the inner part of this disk formed a moon one thousand times the size of Phobos, which has since disappeared. The gravitational interactions created in the outer disk by this massive star apparently acted as a catalyst for the gathering of debris to form other smaller, more distant moons. After a few thousand years, Mars was surrounded by a group of approximately ten small moons and one enormous moon. A few million years later, once the debris disk had dissipated, the tidal effects of Mars brought most of these satellites back down onto the planet, including the very large moon. Only the two most distant small moons, Phobos and Deimos, remained (see the graphic at the end of the press release).

Due to the diversity of physical phenomena involved, no digital simulation is able to modelize the entire process. Pascal Rosenblatt and Sébastien Charnoz's team thus had to combine three successive cutting-edge simulations in order to provide an account of the physics behind the giant collision, the dynamics of the debris resulting from the impact and its accretion to form satellites, and the long-term evolution of these satellites.

In a second study, researchers from the Laboratoire d'astrophysique de Marseille (CNRS/Aix-Marseille Université) ruled out the possibility of a capture on the grounds of statistical arguments based on the compositional diversity of the asteroid belt. They moreover show that the light signature emitted by Phobos and Deimos is incompatible with that of the primordial matter that formed Mars (meteorites such as ordinary chondrite, enstatite chondrite and/or angrite). They therefore support the collision scenario. From this light signature they deduced that the satellites are made of fine-grained dust (smaller than a micrometer³).

Yet the very small size of grains on the surface of Phobos and Deimos cannot, according to the researchers, be solely explained as the consequence of erosion from bombardment by interplanetary dust. This means that the satellites were from the beginning made up of very fine grains, which can only form by gas condensation in the outer area of the debris disk (and not from the magma present in the inner part). Both studies are in agreement on this point. Moreover, the formation of Martian moons from these very fine grains could also be responsible for a high internal porosity, which would explain their surprisingly low density.

The theory of the giant collision, which is corroborated by these two independent studies, could explain why the northern hemisphere of Mars has a lower altitude than the southern hemisphere: the Borealis basin is most probably the remains of a giant collision, such as the one that *in fine* gave birth to Phobos and Deimos. It also helps explain why Mars has two satellites instead of a single one like our Moon, which was also created by a giant collision. This research suggests that the satellite systems that were created depended on the planet's rotational velocity, because at the time Earth was rotating very quickly (in less than four hours), whereas Mars turned six times more slowly.

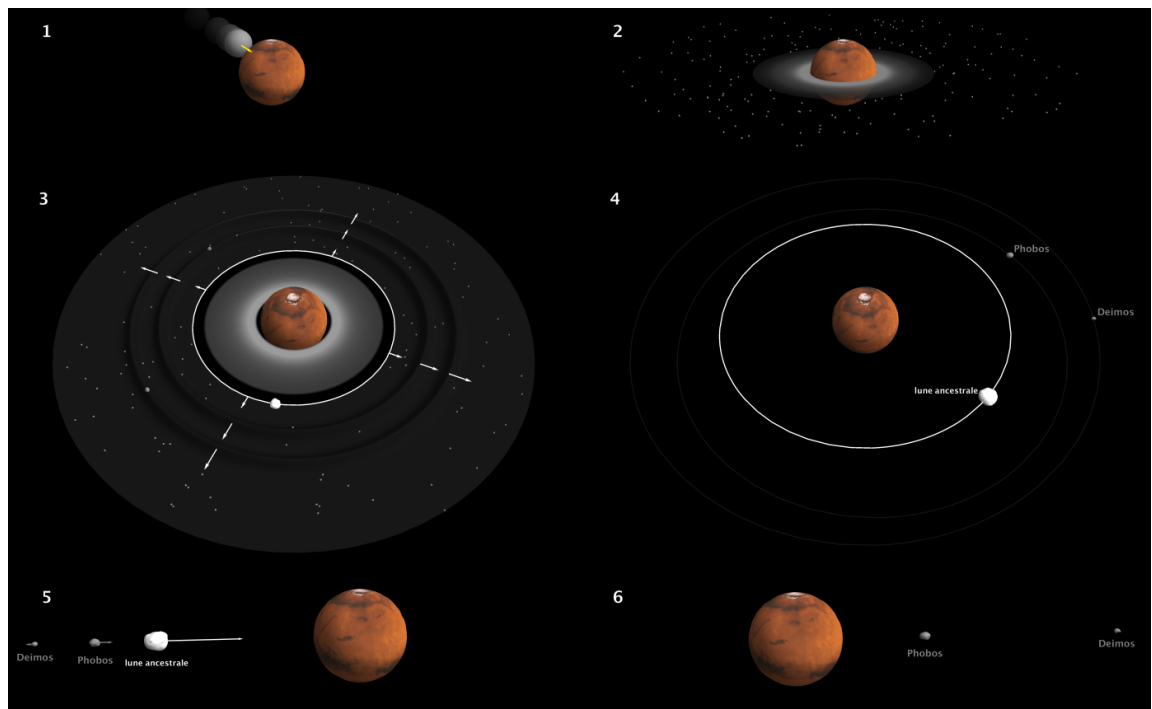
New observations will soon make it possible to know more about the age and composition of Martian moons. Japan's space agency (JAXA) has decided to launch a mission in 2022, named Mars Moons Exploration (MMX), which will bring back samples from Phobos to Earth in 2027. Their analysis could

³ A thousandth of a millimeter.



confirm or invalidate this scenario. The European Space Agency (ESA) has planned a similar mission in 2024 in association with the Russian space agency (Roscosmos).

This research received support from IPGP, the Labex UnivEarthS, ELSI, Kobe University, the Royal Observatory of Belgium and IDEX A*MIDEX.



Chronology of events that may have created Phobos and Deimos.

Mars is struck by a protoplanet one-third its size (1). A debris disk forms within a few hours. The elementary building blocks of Phobos and Deimos (grains smaller than a micrometer) condense directly from gas in the outer part of the disk (2).

The debris disk soon produces a moon near Mars that moves further away and propagates its two areas of dynamical influence like ripples (3), which over the course of a few thousand years causes the accretion of more dispersed debris into two small moons, Phobos and Deimos (4).

Under the effect of the tidal pull of Mars, the large moon falls back to the planet within approximately five million years (5), while smaller Phobos and Deimos take up their current positions in the ensuing billions of years (6).



www.cnrs.fr

université
**PARIS
DIDEROT**
PARIS 7

UNIVERSITÉ DE
RENNES 1

Aix*Marseille
université



Artist's rendering of the giant collision that may have produced Phobos and Deimos along with the Borealis basin.

The colliding body would have been approximately one-third the size of Mars. At the time, Mars was young, and perhaps had a thicker atmosphere and liquid water on its surface.

© Université Paris Diderot / Labex UnivEarthS

To find out more:

An animation film is available: <http://univearths2.in2p3.fr/fr/2016/06/28/phobos-deimos/> (credit: Université Paris Diderot / Labex UnivEarthS)

An article from CNRS News: <https://news.cnrs.fr/node/978>.

Bibliography

Accretion of Phobos and Deimos in an extended debris disc stirred by transient moons, Pascal Rosenblatt, Sébastien Charnoz, Kevin M. Dunseath, Mariko Terao-Dunseath, Antony Trinh, Ryuki Hyodo, Hidenori Genda & Stéven Toupin. *Nature Geoscience*, 4 July 2016. DOI: 10.1038/ngeo2742

Reconciling the orbital and physical properties of the martian moons, Thomas Ronnet, Pierre Vernazza, Olivier Mousis, Bastien Brugger, Pierre Beck, Bertrand Devouard, Olivier Witasse, Fabrice Cipriani. *The Astrophysical Journal*, forthcoming.

Contacts

Université Paris Diderot Researcher | Sébastien Charnoz | T +33 (0)1 83 95 78 29 | charnoz@cea.fr
CNRS Researcher | Pierre Vernazza | T +33 (0)4 91 05 59 11 (office) | pierre.vernazza@lam.fr
CNRS Press Officer | Véronique Etienne | T +33 (0)1 44 96 51 37 | veronique.etienne@cnrs-dir.fr